

## Earth Space Mission and Gravity: Mission to Mars

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### Anchoring phenomenon:

Students will use their understanding of gravity, scale properties, and force to describe the cyclical pattern of movement among Earth, Sun, Moon, and Mars to determine how and when would be a good time to launch a mission to the Moon and later to Mars.

# **Essential question about phenomenon/for unit:** When is the best time to launch a mission to the moon and later to Mars?

### **Big Idea:**

The gravity of the Sun, Earth, and Mars results in attraction between the objects, but since the Sun is so much larger than the Earth and Mars,, the Earth revolves around the Sun every 365.5 days and Mars revolves every 687 days. There is also a gravitational force that causes the Moon to revolve around the Earth approximately every 28 years, and the Moon revolves around the Earth because it is smaller and much closer to Earth than the Sun. In order to show the difference in distance to the Sun, we used a scale model, but had to focus on distance rather than the size of the objects because we would have to have much larger distances to accommodate the size of the objects. There is an ideal time to launch to the moon based on when the Moon is closest to the Earth, but we also have to take into account the idea time to launch from the Moon to Mars. We have to consider when these are aligned.

### Scientific explanation:

Since the Sun is so much more massive than the Earth and Mars, both Earth and Mars revolve around the Earth. There is a gravitational pull between the Earth and Sun, and they are attracted to each other. The Earth pulls on the Sun, but it only moves a little, more like wobbles. The Sun and Mars also have gravity that attracts them to each other, but again the more massive Sun results in Mars revolving around the Sun rather than the Sun revolving around Mars or Earth. With increased massive of the Sun, the gravitational pull would be even greater, which would mean both Earth and Mars would be pulled in closer to the Sun and they would revolve faster. It'd be similar to them being closer to the center of a coin well. In a centripetal force lab, when the mass is increased in order for the object to spin around the same amount, there has to be a greater force from the person. This shows that an increase in mass, leads to an increase in force.

The Sun is so much more massive than the planets and even bigger than our moon. In order to show these differences in size, we

have to sacrifice the accuracy in showing the scaled distances between the objects. Another factor in determining gravitational force is distance. The moon is smaller than the Earth, and definitely smaller than the Sun, so why doesn't the moon revolve around the Sun rather than it revolving around the Earth every 29 days. This happens because it is so much closer to Earth, so the close distance makes more of a difference compared to the size of the objects. So the Earth has a greater gravitational pull on the Moon than the Sun. Additionally gravity explains why Mars and Earth have different orbital speeds and take different amounts of time to orbit around the Sun. Mars is further away from the Sun and it's less massive, so there is less gravitational pull between the Sun and Mars compared to the Sun and Earth. The result of less gravitational pull is a slower orbital speed and a bigger orbit around the Sun, thus taking almost twice as long to revolve around the Sun.

The Earth revolves around the Sun every 365.25 days and Mars revolves around the Sun every 676 days. Since it takes about twice as long for Mars to revolve around the Sun compared to Earth, then theoretically they would be aligned about every 2 years, so the second revolution of Earth around the Sun. However the orbital planes of the planets also have to be aligned and that does not happen every 2 years, so a true alignment of the Sun-Earth-Mars and their orbital planes happens even less than 2 years. Additionally to also have the Sun-Earth-Moon-Mars all aligned is even more rare.

To plan a mission to Mars where we landed on the Moon and then took off to land on Mars would have to be timed perfectly, at best every 2 years. Fortunately our technology does not require us to have orbital plane alignment though. But assuming this perfect alignment was necessary, when the moon was on the other side of the Earth (and Sun was on the other side) that is a full moon. From the human perspective we would be able to see the ½ of the moon that is lit, so it would appear very bright and a full circle. After 14 days the moon will be on the other side of the Earth, and we could not see the ½ that is illuminated, so it is a new moon. It takes about 3 days to travel to the moon, and assuming you wanted to be on the moon for a short period of time perhaps 3-4 days then you should leave the Earth at a first quarter moon, you'd like on the moon at a waxing gibbous. Then you'd stay on the moon until it was a full moon and then leave the moon for Mars.

Standard	РЕ	DCI	ССС
<u>MS-PSS2-4</u>	Construct and present arguments using evidence to support the claim that gravitational interactions are attractive and depend on the masses of interacting objects.	PS2.B Types of Interactions	Systems and System Models
<u>MS-ESS1-3</u>	Analyze and interpret data to determine scale properties of objects in the solar system.	ESS1.A The Universe and Its Stars	Scale, Proportion, and Quantity
<u>MS-ESS1-2</u>	Develop and use a model to describe the role of gravity in the motions within galaxies and the solar system.	ESS1.B Earth and the Solar System	Systems and System Models
<u>MS-ESS1-1</u>	Develop and use a model of the Earth-sun-moon system to describe the cyclic patterns of lunar phases, eclipses of the sun and moon, and seasons.		Patterns
<u>MS-ETS1-1</u>	Define the criteria and constraints of a design problem with sufficient precision to ensure a successful solution, taking into account relevant scientific principles and potential impacts on people and the natural environment that may limit possible solutions.	ETS1.A: Defining and Delimiting Engineering Problems	Influence of Science, Engineering, and Technology on Society and the Natural World All hum

#### NGSS Performance Expectations addressed in this unit:

### Summary Table of Activities in Unit

Activity	<b>Learning Target</b> (Color code 3-dimensions: Blue = SEP; Orange = DCI, Green = CCC)	Evidence Students Could Gain/Key Learning	Connection to explain part of the phenomena
Activity 1: Initial Model	Students will create an initial model that shows the motion of Mars, Moon, and the Earth to explain their proposed timeline to leave for Mars and explain why we'd want to do that.	Objects move around each other. One object makes the others move around it. There is a specific amount of time for these movements that are predictable.	The Moon predictably moves around the Earth and both the Earth and Mars move predictably around the Sun, so there is a supposed ideal time to launch to go to the Moon to later go to Mars.
Activity 2: Criteria and Constraints of the Mission ETS1-1 Exit Ticket: What are the criteria for the mission and justify which are most important. What are the most influential limitations and why?	Ss brainstorm criteria and constraints for different elements of the mission by learning more about the mission.	Going out into space requires understanding gravity and its role in the movement of objects. The more we define the criteria the more successful we can be.	We need to know the distance to get to the Moon and Mars and the velocity of their orbits as well as the velocity of our spacecraft. Additionally we need to know the environment so we can protect our bodies in space.
Activity 3: Mission #1 - Travel to Mars (pgs. 19-24) ESS1-3 Exit Ticket: How does the distance to the Sun compare from the Earth to Mars? What do you think the Sun looks like on Mars compared to Earth? Scale Properties for Student Scale Properties KEY	Students study interplanetary distances and scale to revise their models and predict when to launch.	Mars is 1.5x further away from the Sun compared to Earth, resulting in it moving slower around the Earth and it takes almost twice as long to orbit the Earth. A scaled model requires choosing between distance or diameter of planets and each has a strength and weakness.	Mars is further away from the Sun, so the model shows that it is has a bigger radius and orbit around the Sun compared to Earth.

Activity 4: Build solar sails to explore how propulsion is used for spacecraft. (Optional - cont'd of previous lesson, starting on page 6) ETS1-1	Students will construct a scale model solar sail and experience the tradeoffs, criteria, and constraints.	Students will research and build a scale model of a solar sail and habitat pod. Students will discuss the tradeoffs of their model as well as the criteria and constraints.	When engineering a product, there are criteria and constraints that you must work with. Additionally there are tradeoffs, like size vs stability or size vs cost that have to be considered.
Activity 5: Gravity and Orbital Motion ESS1-2 <u>Reading Gravity and Orbits</u> <u>Phet simulation</u>	Students will determine how distance and mass affect gravity resulting in a difference in orbital motion in a simulation model.	Gravity is an attractive force that keeps planets in orbit around a more massive object. As the distance increases the force of gravity is less, so the planet travels slower and in a larger orbit. Planets that are closer to the Sun and more massive have a greater gravitational pull, so they move faster around the Sun and in a smaller orbit. The phet simulation shows that the moon also orbits around the Earth because it have a greater gravitational pull towards the Earth since it is so close.	Mars is further away from the Sun and less massive, so it moves slower and has a larger orbit around the Sun compared to the Earth. Earth is closer and more massive, so it moves faster and has a smaller orbit around the Sun. They're both attracted to the Sun because of gravity and therefore orbit around the Sun.
Activity 6: Moon Phases <sup>ESS1-1</sup>	Students will explain how the movement of the moon and our perspective results in cyclical moon phases we see.	The moon is always half lit up, but as it travels around the Earth approximately every 29 days, our perspective changes. Each day we see a different part of the lit up moon as it travels. When the moon is between the Earth and Sun, we do not see the lit side, so we only see the dark full moon. But after 7 days of travel, it is now at	It takes us approximately 3 days to travel to the moon. Depending on how long you want to spend on the moon (in my model approximately 3-4 days), you could leave the Earth when the moon was a first quarter moon, but you'd arrive at the moon when it's a waxing gibbous. You'd stay on the moon for 3-4 more days and leave when it was aligned as the

		a 90 degree angle, and we see a first quarter moon. Seven days later, the moon has traveled so the Earth is between the moon and the Sun, and we see the full lit up side of the moon, which is a Full moon.	Sun-Earth-Moon, so it'd appear as a full moon from the Earth.
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## Model Template

When and how can we go to Mars but land on the Moon first?			
In your initial model, include:  A picture that shows the location, movement, and speed of Sun, Earth, Moon, and Mars factors that might affect the MOTION of these objects in space relative sizes and distances of the objects from each other suggestion of when we launch and rationale for why	SunMoonMarsEarthImage: Sun seriesImage: Sun series		
	I think we should launch to Mars when		

#### **Initial Models**









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#### **Final Models**



#### WATER BALLOD Weaknesses + THE SUN WERE BIGGER. The orbit of the Earth and Mors will be faster ? The water balloon lab contradicled The water ballion tab consections the Galt theil an increase in Galtance decreases the gravitation Galtance decreases the gravitation fine longer stilling (SAI) product a factor underly than the SAI string. This way, to change the this, we need to change the I know this because in the PHET simulation -When I increased the mass of the Sun, the gausticianal forced also increased. Thus, Maro and Barth will be pulled door resulting in a smaller what One weakness in my model is it a size of the planets. If the Sun were that big ( Ilinches in diameder. a "Unajer orbit. - Ind because the Mars and Earth were clear their orbital speed increased. I noticed the number of days Earth and Mars tack to orbit the san decreased greatly why? When qualifolding force increases, calculat one of the same Earth should be miniecule. & 1.3million fime less than Mors and the Mannshould Two main problems 1. THE WATER BALLOON be even smaller WAS TOO HEAVY speed increases. In this model, the water \*I also know this because of our solar system right new Comparing Earth and Mars, Mars gres Slower than Earth Ceause it is further oway. If One strength in my model ballion's mass was almost heavier than the center of its orbit (our is its distance scale. Mars is 227 M Km away from the SUN. I scaled my model SD I cm & T.M. Km COO It is reas anable that Mars and Earth go doser, their orbital speed hand). The ballion should be lighter because it was will increase. harder to swing the ballion . on a 5.9 string than 3.1.0 Thus we had to swing failth on the 5.9 string to keep it in the air. I made Mars & 32 cm IF THE SUN WERE SMALLER. a way from the Sun. Everything is in place where it should be Earth is closer . The orbit of the Earth and Mars will be slowers larger. I know this because in the PHET simulation. it in the air. to Mars. The Maan is doser 2. DATA WAS INCONSES--When I decreased the mass of the Sun, the gravitatingl fore also decreased Thus, Mars and to the Earth. And we can still TENT See everything on. G small We had 3 different people Earth will be pulled doser resulting in a larger spin the water ballions. I for one had a tough piece of paper. Moon Further Explanation - And because the Mors and Earth whe closer, their Moon to the Man shifts phase. And it is Even 3 days, the Man shifts phase. And it also takes is 3 day to get to the Maon (Protet yes is 3333 mph) We need 3 days real on the Man time w/ the 5 foot string into active remained and paint recent, we ordered generated . I relieve the number of day section and place the to arbit the san increased grading why? When granitational fire decreases, which speed decreases. while others had an easter time spinning 14. I subsequently swung my 5 feet sking faster so So, if we leave on a First Quarter Moon () 20, It we note at a trust countertaint (). We will get the on a Washs Count (). After 3 days real, the Maan will be perfectly. In Delivern Mars and Eath (a Rui Maan ()) I also know this because of our solar system right now. the balloon wouldn't pop. Instead of 3 people 1 person would be better. Comparing Earth and Mors, Mars goes sharer than Earth because it to Europe above, I Chang and Earth go Further, or closest to Mars. And then we can goto their unbital speed will degrade. Mark

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### **Final Model Rubric**

Mission to Mars Modeling Rubric				
PSS2-4 Construct and present argume of interacting objects.	nts using evidence to support the claim that gravitationa	al interactions are attractive and depend on the masses		
Approaching (2)	Meeting (3.6)	Exceeding (4)		
<ul> <li>Describes the phet simulation, but does not explain any patterns we saw.</li> </ul>	<ul> <li>Explains how the force of gravity changes when the mass of objects change.</li> <li>Describes how we saw those changes in the phet simulation and how that helps explain how Earth and Mars are different (need to find their masses).</li> </ul>	<ul> <li>Explains how we could change the water balloon lab to try to model this.</li> <li>Explains how changing the mass of the Sun would change your entire model.</li> </ul>		
ESS1-1 Develop and use a model of the Earth-sun-moon system to describe the cyclic patterns of lunar phases, eclipses of the sun and moon, and seasons.				
Approaching (2)	Meeting (3.6)	Exceeding (4)		
Draws where the moon would be positioned when we leave for, land on, and leave from the moon and determines how long it would take.	<ul> <li>Shows where the moon would be when when we leave for, land on, and leave from the moon.</li> <li>Draws and explains the moon phases during those 3 times.</li> </ul>	Justifies why we would leave Earth and Moon based on how long it takes to travel in the rocket and the distance the moon is from the Earth.		
ESS1-2 Develop and use a model to describe the role of gravity in the motions within galaxies and the solar system.				
Approaching (2)	Meeting (3.6)	Exceeding (4)		
Makes basic "If, then" statements about the role of gravity in the motion in our solar system.	<ul> <li>Shows and explains bigger and smaller gravitational pull</li> <li>Shows the "balance" between gravity and inertia that results in orbital motion</li> <li>Relates orbital motion (speed) to gravitational pull.</li> </ul>	<ul> <li>Calculates the force of gravity for the different planets and the moon</li> <li>Explains how the gravitational pull and motion of planets would change if Sun became bigger or smaller.</li> <li>Shows these changes if Sun were bigger or smaller.</li> <li>Uses evidence either from demonstrations or data to support ideas.</li> </ul>		
ESS1-3 Analyze and interpret data to determine scale properties of objects in the solar system.				
Approaching (2)	Meeting (3.6)	Exceeding (4)		
Communicates data about either the distance, mass, or velocity of objects in the solar system.	<ul> <li>Model is scaled either with the size of the objects or their distance.</li> <li>Explains how Earth and Mars are different</li> </ul>	<ul> <li>Explains why you chose to prioritize distance or size.</li> <li>Describes a weakness and a strength of your model.</li> </ul>		

### Additional Unit-Related Documents

Assessment on Google Form